IRIS.

Ephemeris at her Reappearance, 1848. By Mr. Hind.

"The following ephemeris of this planet up to November 11, is founded upon elements calculated from the observations at Mr. Bishop's Observatory on August 13, 1847; Dr. Wichmann's with the Königsberg Heliometer, on November 7; and Professor Challis' latest observation, 1848, February 17. The observed positions were corrected for aberration and parallax; but the effect of planetary perturbations on the geocentric place during the interval, being very minute, has been neglected.

Epoch 1847, August 1.0, Greenwich Mean Time.

Mean Longitude of	<i>Iri</i>	s	33 I	24 24	20.57	M 12
	7		41	31	19.71	(Aug. 1.
	\mathcal{B}		2 59	44	45'91	
	φ	••••••	13	19	59.98	
\mathbf{Log}	æ				0.3772	:589
	μ	••••••		91	54"•091	333

"These elements represent the middle declination exactly, and give the right ascension too small by o"8. On comparing them with the latest observation which I have yet seen, viz. Professor Kaiser's on March 14, (published in a recent number of the Society's Monthly Notices), I find the following errors:—

Cos.
$$\delta \times \Delta \alpha = -2'' \cdot 37$$
 $\Delta \delta = -5'' \cdot 11$

"These differences being so very small, it appeared to me probable that any attempt to correct the orbit from a combination of the observations during the first apparition of *Iris* only, would hardly repay the trouble that must necessarily be expended upon it. I have therefore delayed the final investigation until after the next appearance of the planet. The two oppositions can then be connected, and the elements resulting from a discussion of all the observations will no doubt be sufficiently exact to form the basis for ephemerides for some time to come.

"Assuming, therefore, that the orbit just given applies to Greenwich mean noon of 1847, August 1, I have calculated the perturbations produced by Venus, the Earth, Mars, Jupiter, and Saturn, between this date and 1849, February 3, (nearly the epoch of opposition), employing the same values for the masses as are now adopted by Encke in his investigations on the comet which bears his name; their logarithms are,—

Venus	4.39595
Earth and Moon	4.44916
Mars	3.5718
Jupiter	6.97969
Saturn	6.45573

"The united effects of these planets, between 1847, August 1, and 1849, February 3, are,—

$$\int d\alpha = +0.00053771$$

$$\int de = +0.000302365$$

$$\int d\alpha = +295''.282$$

$$\int dL = +146.486$$

$$\int d\Omega = -76.789$$

$$\int di = -7.232$$

"For the ephemeris subjoined the true elements were obtained for every 24th day, and interpolated with 4th differences, where sensible, for every 8th day. The geocentric places so obtained were interpolated with 4th differences for each day. For the sake of brevity the elements of the variable ellipse are here given for the extreme dates of the ephemeris only:—

		Aug. 7 ^d ·0	Nov. 11 ^d .0
-	M	29 28 26.64	0 / // 55 9 7°09
	W.	41 35 29.24	41 36 55.13
	8	259 45 41.02	259 45 51.44
	i	5 28 14.30	5 28 11.39
	φ	13 20 0.52	13 20 33.03
Log.	pe	2.9841593	2.9840273

"The longitudes are referred to the apparent equinox of date; corrections being applied for diminution of obliquity, &c. The ephemeris contains the position of the planet, for the apparent equinox and equator of each date, unaffected with aberration; consequently, before comparing with apparent places, it will be necessary to subtract the time given in the last column from the mean time of observation.

Ephemeris.

At Greenwich Mean Noon.

-0.0		R.A.	Decl.	Log. Δ	497°·8 × △
1848. August	7	0 / //		8 0:42704	22 44·6
Tragast	7	104 51 34			22 44 0
	8	105 27 11.	3 22 6 11.3	3	
	9	106 2 40	22 0 57.2	2 0.43679	22 40.9
	10	106 38 2	0 21 55 35	4	
	11	107 13 15	9 21 50 6.	3 0.43560	22 37.2
	12	107 48 22	0 21 44 29"	7	
	13	108 23 20	21 38 45	9 0.43437	22 33.4
	14	108 58 10	•5 21 32 54.8	3	
	15	109 32 52	·8 21 26 56·7	0.43310	22 29.4
	16	110 7 27	21 20 51.6	5	
	17	110 41 53	3 21 14 39.6	6 0*43178	22 25.3
	18	111 16 11	4 +21 8 20.8	3	

-0.0		R.A.	Decl.	Log. Δ	$497^{\text{s} \cdot \text{8}} \times \Delta$
1848. August	19	111 50 21.0	+21 1 55.4	0.43042	22 21·I
_	20	112 24 22.3	20 55 23.3		
	2 Ï	112 58 15.1	20 48 44.8	0.42900	22 16.7
	22	113 31 59.3	20 41 59.8		
	23	114 5 35.0	20 35 8.6	0.42755	22 12.3
	24	114 39 2.0	20 28 11.1		
	25	115 12 20.2	20 2r 7.4	0.42604	22 7.7
	26	115 45 29.6	20 13 57.8	_	
	27	116 18 29.9	20 6 42.2	0.42448	22 2.9
	28	116 51 21.5	19 59 20.9	00	a
	29	117 24 3.2	19 51 53.9	0.42288	21 58.0
	30	117 56 36.1	19 44 21.3	0140700	21 53.0
C4	31	118 28 59.6	19 36 43.3	0.42122	21 55 0
Sept.	. 1	119 33 18.2	19 21 11.3 19 28 59.9	0.41921	21 47'9
	2	119 33 18 2	19 13 17.6	o 4-33-	47 /
	3 4	120 36 58.5	19 5 18.8	0.41772	21 42.6
	5	121 8 34.3	18 57 15.0	4.773	•
	6	121 40 0.3	18 49 6.5	0.41594	21 37.2
•	7 .	122 11 16.6	18 40 53.2		
	8	122 42 23.1	18 32 35.3	0.41407	21 31.6
	9	123 13 19.7	18 24 12.9		
	10	123 44 6.5	18 15 46.1	0.41215	21 25.6
	11	124 14 43.4	18 7 15.0		
	12	124 45 10.3	17 58 39.7	0.41012	21 200
	13	125 15 27.2	17 50 0.3		
	14	125 45 34.0	17 41 17.0	0.40814	21 14.0
	15	126 15 30.8	17 32 29.7	_	
	16	126 45 17.4	17 23 38.7	0.40602	21 7.9
	17	127 14 54.0	17 14 44.0		
	18	127 44 20.3		0.40390	21 1.7
	19	128 42 41.8 128 42 41.8		0.40170	20 55.3
	20 21	129 42 41 8		0 401/0	20 33 3
	21	129 40 21.3	_	0.39944	20 48.8
	23	130 8 22.0	_	- 37714	4
	-3 24	130 37 17.9		0.39712	20 42'1
	25	131 2 30.1	•	37,	·
	26	131 33 31.5		0.39473	20 35.3
	27		15 42 40.3		· · · · ·
	28	132 29 0.2	_	0.39229	20 28.4
	29	132 56 28.3	15 24 42.9		
	30	133 23 44.9	15 14 10.9	0.38978	20 21.3
O	ct. I	133 50 50.5	+ 15 4 36.9		

Iris.

184	8.	R.A.	. ,,	0	Dec	el. "	Log. 🛆	497 ^s ·8 × Δ m s
Oct.	2	134 17				o · 9	0.38720	
	3	134 44	26.3	14	45	23.5		
	4	135 10	56.9	14	35	43.7	0.38456	20 6.7
	5	135 37	16.0	14	26	2.6		
	6	136 3	23.5	14	. 16	20° I	0.38184	19 59.7
	7	136 29	18.8	14	6	36.1		
	8	136 55	2.6	13	56	50.9	0.37910	19 51.7
	9	137 20	34.5	13	47	4.2		
	10	137 45	54.4	13	37	17.0	0.37628	19 43.9
	11	138 11	2.4	13	27	28.5	•	
	12	138 35	58.3	13	17	39.5	0.37339	19 36.1
	13	139 0	42° I	13	7	49.2		
	14	139 25	13.6	12	57	58.6	0.37044	19 28.1
	15	139 49	32.8	12	48	7°5		•
	16	140 13	39.2	12	38	16.1	0.36741	19 20'0
	17	141 37	33.6	12	28	24.4		•
	18	141 1	15.0	12	18	32.2	0.36432	19 11.8
	19	141 24	43.6	I 2	8	40.2		
	20	141 47	59°4	11	58	48.6	0.36112	19 3.2
	2 I	142 11	2.0	11	48	56.9		
	22	142 33	51.4	11	39	5.2	0*35795	18 55.0
	23	142 56	27.4	11	29	14.7		
	24	143 18	50.0	11	19	24.3	0.35465	18 46.4
	25	143 40	58.9	11	9	34.9		
	26	144 2	54.0	10	59	46.1	0.32129	18 37.7
	27	144 24	35.1	10	49	58.4	:	
	28	144 46	2.0	10	40	11.8	0.34785	18 28.9
	29	145 7	14.6	10	30	26.5		-
	30	145 28	12.7	10	20	42.2	0.34436	18 20.0
	31	145 48	56.3	το	11	0,0		
Nov.	I	146 9					0.34079	18 11.0
	2	146 29	39.1	9	5 I	40.1		
	3	146 49	38.1	9	42	2.8	0.33716	18 2.0
	4	147 9	22.0	9	32	27.5		
	5	147 28	50.2	9	22	54.2	0.33346	18 52.8
	6	147 48			13	23°2		
	7	148 7	-	-		54.5		17 43.6
	8	148 25						
	9	148 44			45	4.9	0.32584	17 34.2
	10	149 2	19.5	8	3,5	44.3		
	11	149 20	13.1	+8	26	26.4	0.32102	17 24.8